## X-rays and high pressure

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Shockwave experiment **Ricardo Samad (IPEN) Nilson Dias (IPEN) Jefferson Bettini (LNNano) Raul Freitas (LNLS)** 

## X-rays and pressure in the universe: Stars, supernovas, ...

#### **Neutron star from Nasa**



#### Range of pressure in the universe



From: https://www.gl.ciw.edu/static/users/rhemley/HemleyWilliamson2004.htm



#### Pressure to make you hungry...

**Pressure cooking since 1679** (Wikipedia) Lid seals shut trapping steam inside Trapped steam builds pressure and raises cooking temperature. Wet cooking (steaming and boiling) speeds heat transfer to food. P up to 2 atm **Cold press Pasteurization:** Pressed to about 0.5 GPa (5000 atm)

## PV = nRT

A lot of scientific studies on foods under pressure!!!



From Amazon.com



## High pressure crystallography

#### Macromolecular structures investigated under pressure



**Also:** high-pressure cryocooling methods for protein crystals to preserve them and to diffract better.



#### Pressure and temperature inside earth



From: DiscoveryMagazine



#### High pressure crystallography



X-ray diffraction under pressure with the advent of microfocus beamlines

Enormous impact to study earths interior in the last 20 years!!!



#### **Pressure = Force / Area**





4 Ton Elephant on one foot:

**Pressure =** 

65 kg woman on 1cm<sup>2</sup> heel: Pressure =

1 Bar = 0.987 Atmosphere



#### **Pressure = Force / Area**



Pressure = 190 Ton / 1 cm<sup>2</sup> = 19 GPa (1.9x10<sup>5</sup> Bar) (Easily achievable today at the XDS beamline downstairs)



#### **XRD under pressure**









## **XRD at High Pressures @LNLS**

## Up to 80 GPa for now

- High energies: > 25 keV
- Small beam sizes: < 0.15 x 0.15  $mm^2$
- In-situ pressure calibration,
- Temperature control
  - 10 K (cryostat)
  - 3000 K (laser heating, comissioning)











- Volume collapse with pressure: almost all rare earths
- Mixed valence: half of all rare earths
- 4f electrons: magnetism, electronic structure





## **XRD at Sector 16/HPCAT/APS**: structure

- Lattice contraction as a function of pressure
- Electronic instabilities at low pressure
- 1.00 -– B1 B1 -> B2 transition 0.95 - B1 0.770 **B2** 0.90 Local volume expansion 10 11 12 13 0 760 0.85 0 754 Coordination change 0.80 B2: CsCl B1: NaCl 0.75 - Predicted valence 0.70 transition to Eu<sup>3+</sup> 0.65 0.60

10

n

20

30

40

70

60

50 Pressure (GPa) 80

90



Valence change towards  $3 + (4f^6, 5d^1)$  Up to 40 GPa, as expected.

Contrary to the expected valence goes from 2.3+ back to 2+







## **Homogeneous VS inhomogeneous**

- Homogeneous valence at each crystal phase
- Inhomogeneous valence in the coexisting phase region



## (Dis)A

## (Dis)Agreement with theory



#### **Bond-valence parameterization:**

Brese & O'Keeffe, ACTA Cryst. B 47, 192 (1991). **"bond length is a unique function of bond valence"** 

$$v_{ij} = \exp[(R_{ij} - d_{ij})/b].$$
  $\sum_j v_{ij} = V_i.$ 

**Constants:** 

$$\mathbf{b} = 0.37 \text{ Å}$$
;  $\mathbf{R}_{ij}$  (Eu<sup>2+</sup>) = 2.147 Å;  $\mathbf{R}_{ij}$  (Eu<sup>3+</sup>) = 2.076 Å





#### X-ray Magnetic Circular Dichroism





## **High pressure XAS/XMCD**



## 4-ID-D beamline

#### **Advanced Photon Source**

In-situ pressure calibration: DAC translated to a Ruby fluorescence station

Performed experiments up to **90 GPa**, so far. P > 150 GPa are possible

High Pressure Research 28, 185 (2008) Rev. Sci. Instrum., 78, 083904 (2007)





## at APS: Europium monochalcogenides

T<sub>c</sub> increases with lattice contraction: 4f – 5d mixing (indirect exchange)





## **EuO under pressure**

**P < 14 GPa:** quantum state of fractional 4f/5d occupations

**14 < P < 44 GPa:** fluctuating valence between Eu<sup>2+</sup> and Eu<sup>3+</sup>; homogenously distributed; with characteristic frequency determined by the 4f bandwidth.

**44 < P < 59 GPa:** inhomogeneous valences concomitant with the mixture of NaCl and CsCl phases

**P > 59 GPa:** reentrant Eu<sup>2+</sup> valence ground state is unexpectedly recovered!

Bond-valence parameterization does not work universally

Strong implications to the physics of 4f systems

Science editors' choice v337 p504, (2012).





Pressure = 50000 Ton / 1 cm<sup>2</sup> = 5 TPa (5x10<sup>8</sup> Atm) (That's more challenging and motivating)



• Must go for extreme pressure, temperature and field



## EMA beamline at SIRIUS

A3m-SR

## Extreme condition XRD/XMCD/XES/XRS

# With very small x-ray beamsize



## In a good day



## **High pressure with DAC**





### **General specifications**

Source		SCU16.5 or IVU19 @ lo	ow β-straight
Energy range		2.7 keV to 35 keV	
Monochromator		Vertical bounce DCM, S	i (111) & Si (220)
Polarization control		Circular/Linear using ¼ wave plate	
Photon Flux hutch: 10 keV	<ul><li>@ Magnet</li><li>@ Laser hutch:</li></ul>	<ul> <li>1×10<sup>14</sup> photons/s (direct focusing, DF)</li> <li>4x10<sup>13</sup> photons/s (direct focusing, DF)</li> <li>2x10<sup>12</sup> photons/s (secondary source, SS)</li> </ul>	
Beamsize hutch: (divergence) 10 keV	<ul><li>@ Magnet</li><li>@ Laser hutch:</li></ul>	<b>2x1 μm<sup>2</sup></b> ( 1x0.2 mrad <sup>2</sup> 0.5x0.5 μm <sup>2</sup> ( 4x0.4 mra <b>0.3x0.5 μm<sup>2</sup></b> ( 0.7x0.4 μ	) (DF) ad <sup>2</sup> ) (DF) mrad <sup>2</sup> ) (SS)
Flux at sample Direct focusing Laser hutch (for magnet hut	e w/ g at ch, x3) 10 <sup>17</sup> 10 <sup>17</sup> 10 <sup>17</sup> 10 <sup>17</sup> 10 <sup>17</sup> 10 <sup>17</sup> 10 <sup>17</sup> 10 <sup>17</sup> 10 <sup>17</sup> 10 <sup>17</sup>	<b>4 mrad</b> 10 <sup>14</sup> 10 <sup>13</sup> 10 <sup>12</sup> 10 <sup>11</sup> 10 <sup>10</sup> 10 <sup>15</sup> 20 25 30 35 40 E [keV]	2.3 mrad 5 10 15 20 25 30 35 4 E [keV]



## **XMCD/XMLD** under pressure

#### Large bore SC magnet



#### < 2µm<sup>2</sup> focused beamsize With long working distance

## **Custom anvil cells**





## **EMA beamline at SIRIUS**

High pressure Lab

Saser Hutch

LaserLab

Storage

Iagner

User Space

## Even more extreme pressures in the future ???

Ac machine



## **Shockwaves + X-rays**







## Shockwave + Sirius ?

### Laser induced extreme conditions experiments at Sirius



By Francisco Maia et al.

### **Proof of concept shockwave experiment**

Irradiate graphite with high power laser!



#### SCIENTIFIC REPORTS OPEN Synthesis of diamond-like phase from graphite by ultrafast laser driven dynamical compression

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#### **Electron diffraction of nanocrystals**



#### "Nano onion rings":



## Extreme pressure, temperature, magnetic field in the universe

Lots of other high pressure synchrotron techniques

A playground for discovering new materials to improve life and understand the universe.



From: www.nasa.gov

From: Spring8:





**Questions?**